

Production and properties of superheavy elements

V. Yu. DENISOV

Institute for Nuclear Research, Kiev, Ukraine

Ukrainian-German Conference:

Possibilities of cooperation in the framework of the international project

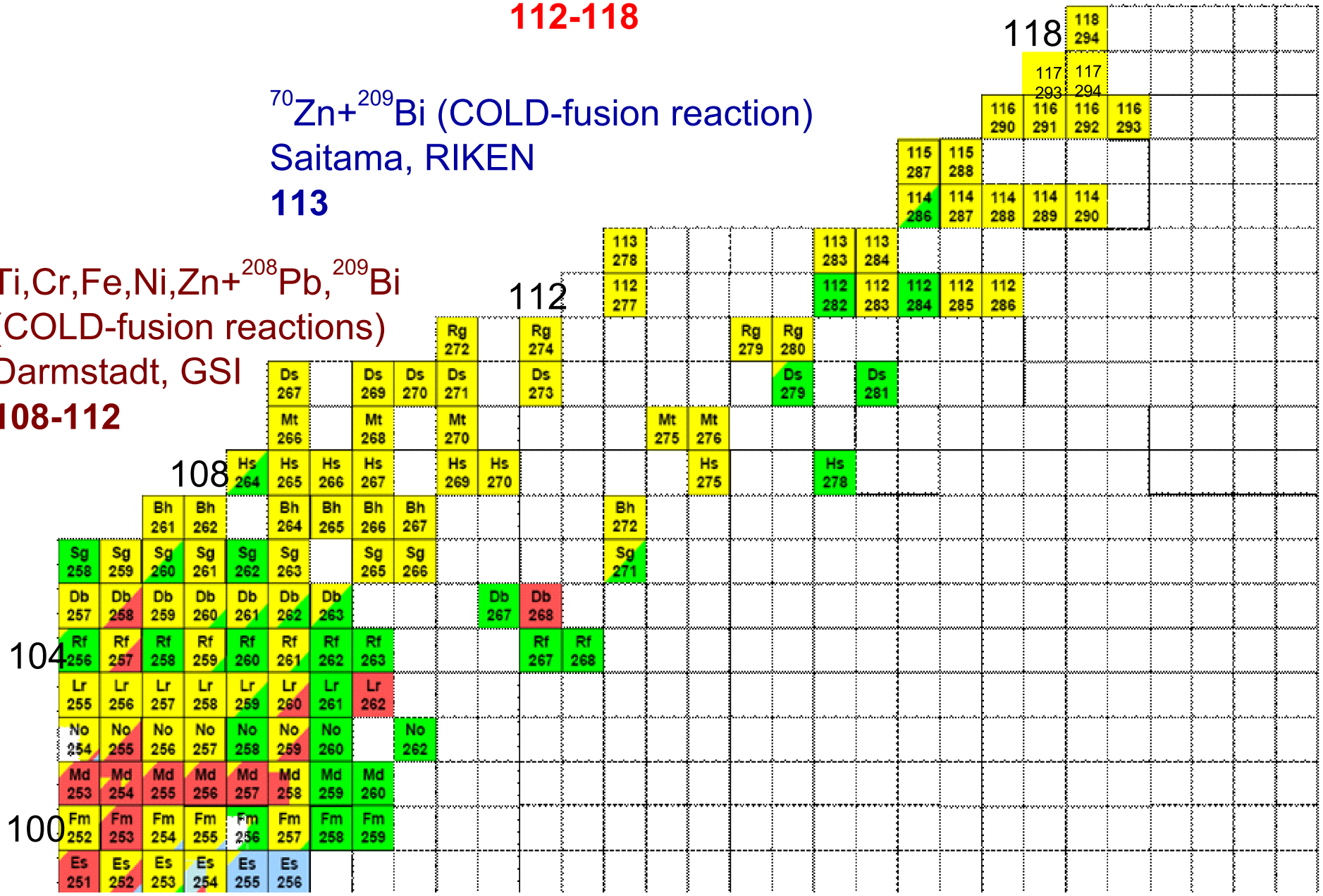
FAIR

September 26, 2014

$^{48}\text{Ca} + \text{U, Pu, Am, Cm, Bk, Cf}$ (HOT-fusion reactions)
 Dubna, JINR
112-118

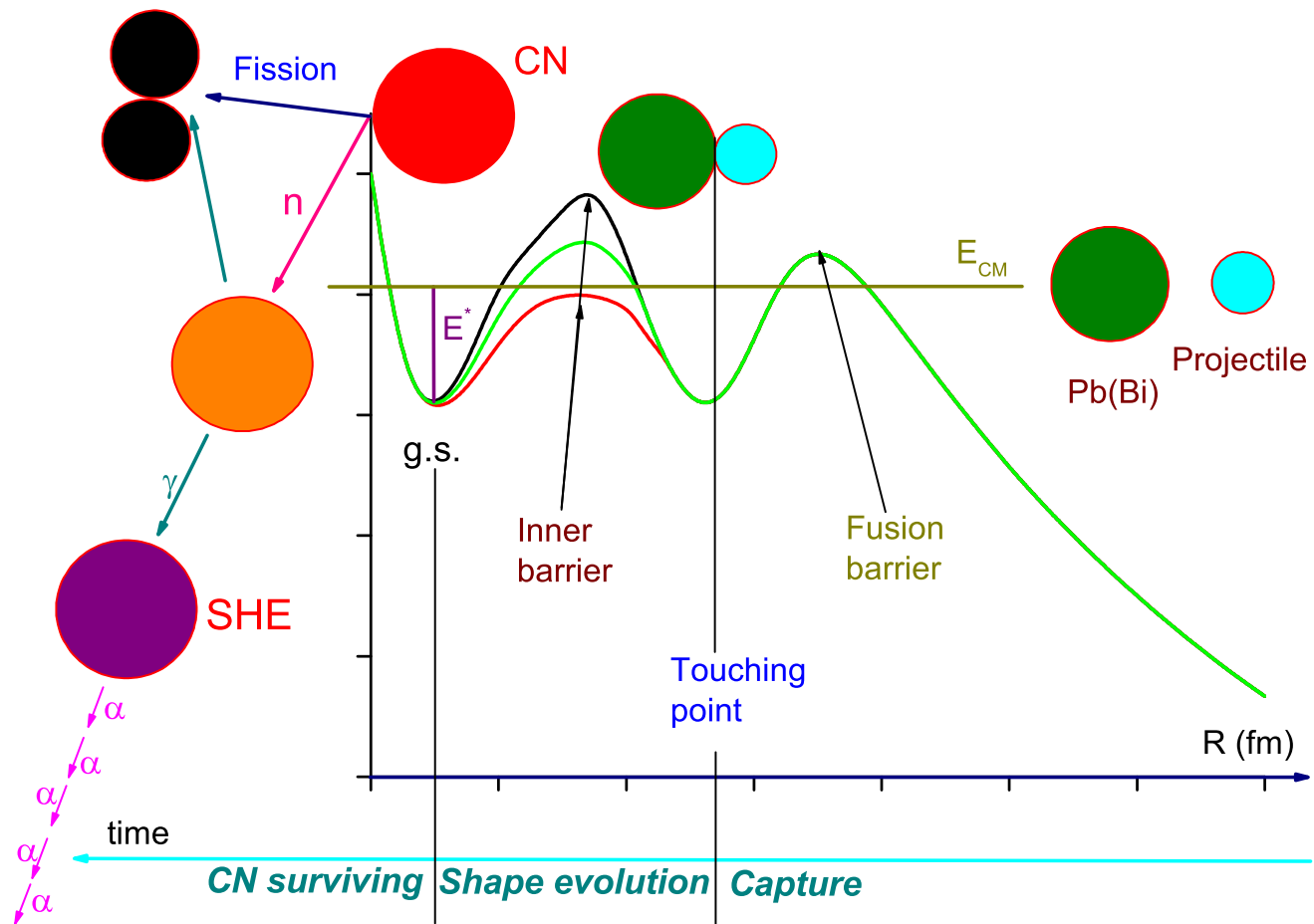
$^{70}\text{Zn} + ^{209}\text{Bi}$ (COLD-fusion reaction)
 Saitama, RIKEN
113

$\text{Ti, Cr, Fe, Ni, Zn} + ^{208}\text{Pb}, ^{209}\text{Bi}$
 (COLD-fusion reactions)
 Darmstadt, GSI
108-112



Main feature of the model

$$\sigma_{\text{SHE}}(E) = \frac{\pi \hbar^2}{2\mu E} \sum_{\ell} (2\ell + 1) \cdot \mathcal{W}_{\text{surv}}(E, \ell) \cdot T_{\text{CN}}(E, B_{\text{inner}}, \ell) \cdot T_{\text{capture}}(E, B_{\text{fusion}}, \ell).$$



Capture

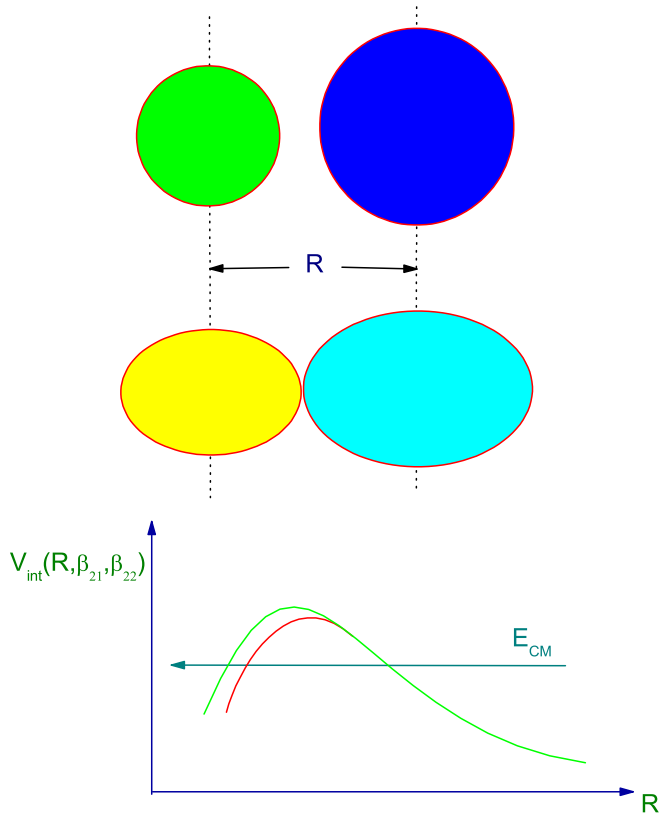
Interaction potential: $V_{\text{Nucleus-Nucleus}}(R, \ell) = V_{\text{nuclear}}(R) + V_{\text{Coulomb}}(R) + \frac{\hbar^2 \ell(\ell+1)}{2MR^2}$

Penetration through barrier: WKB approximation

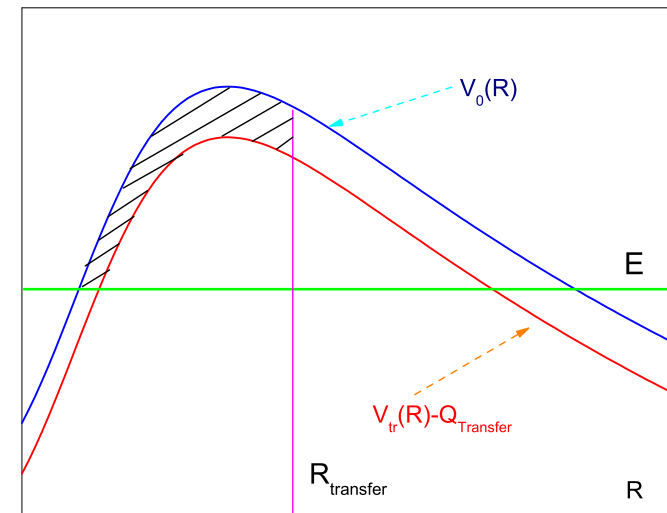
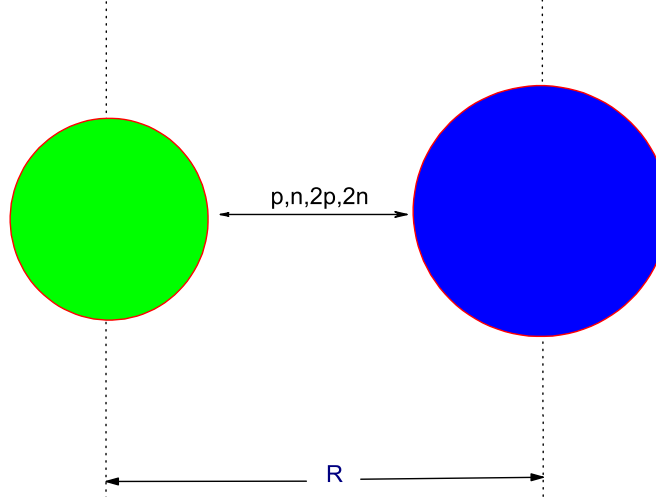
Enhancement of the penetrability due to:

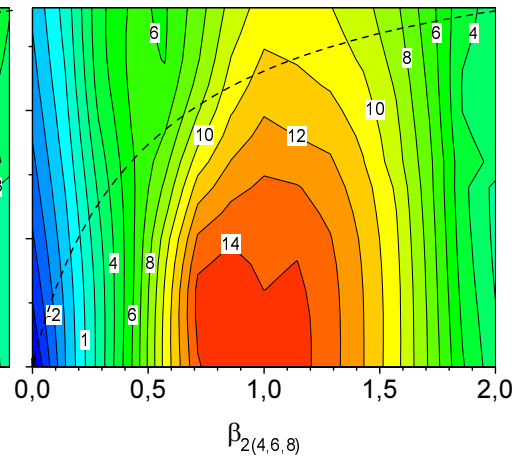
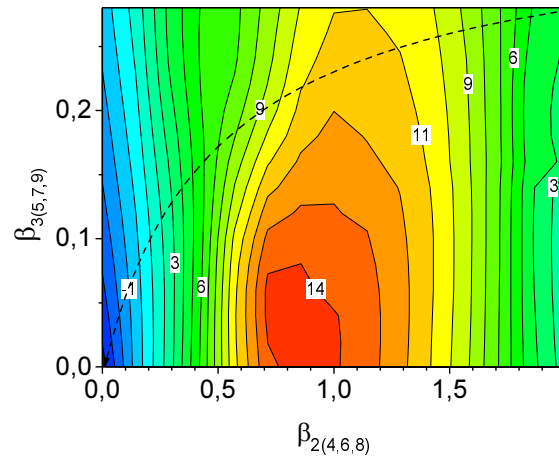
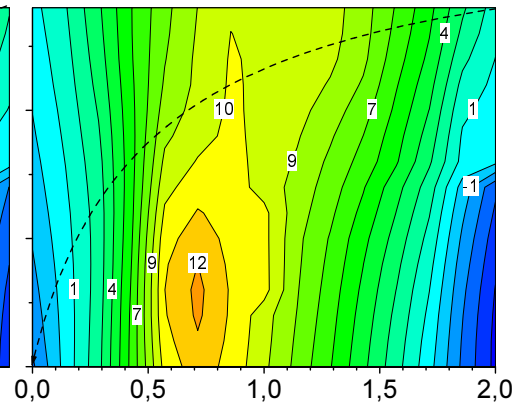
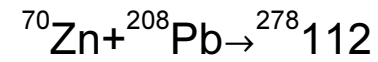
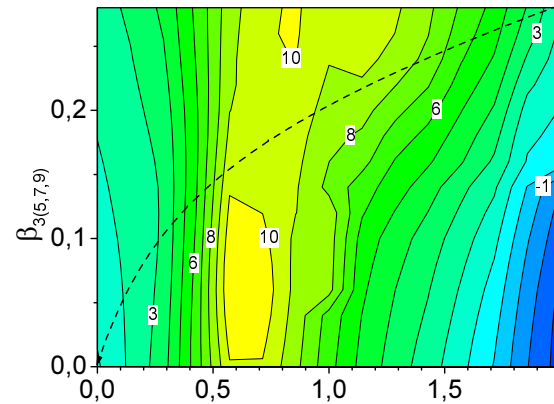
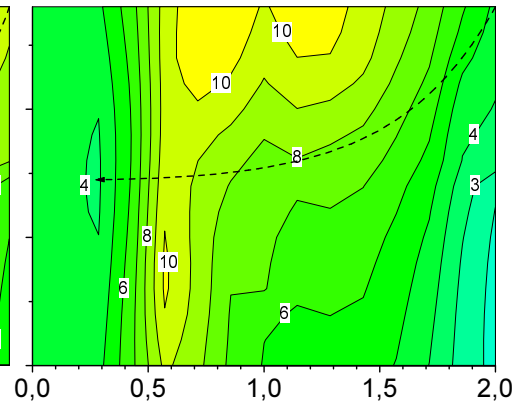
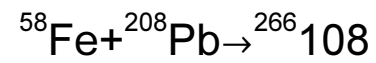
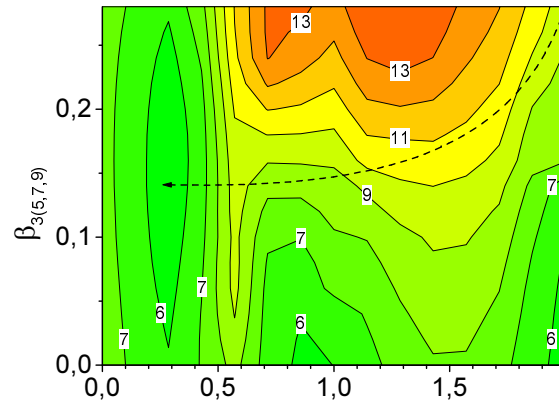
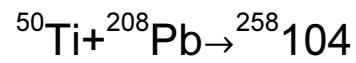
- low-energy surface 2^+ , 3^- vibrations
- nucleon transfer with positive Q -value

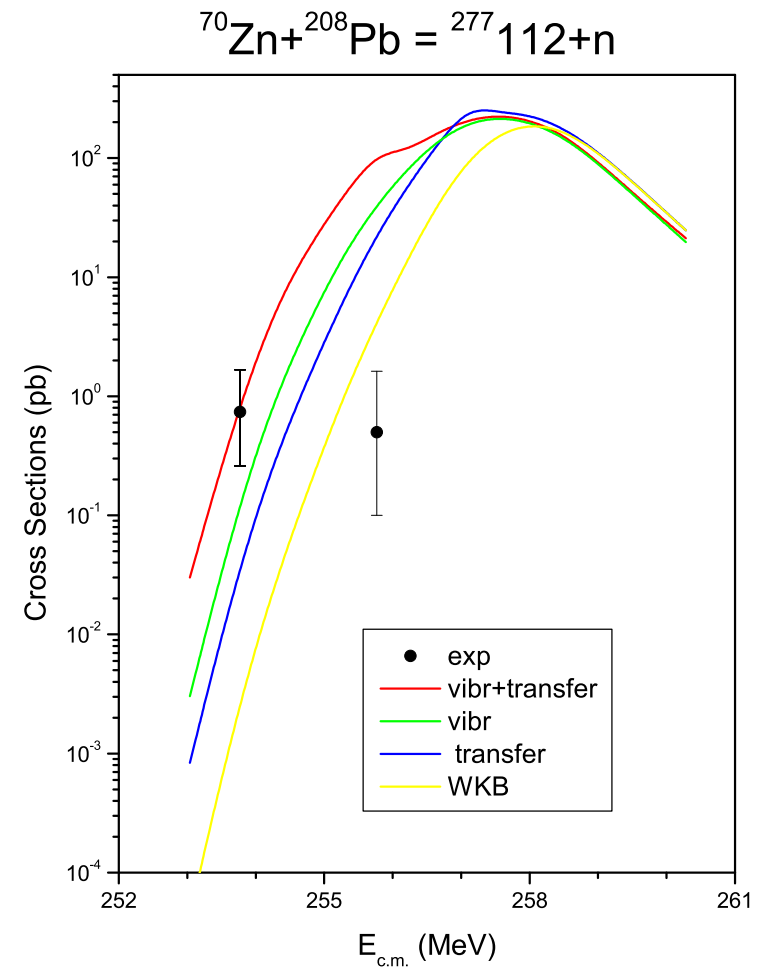
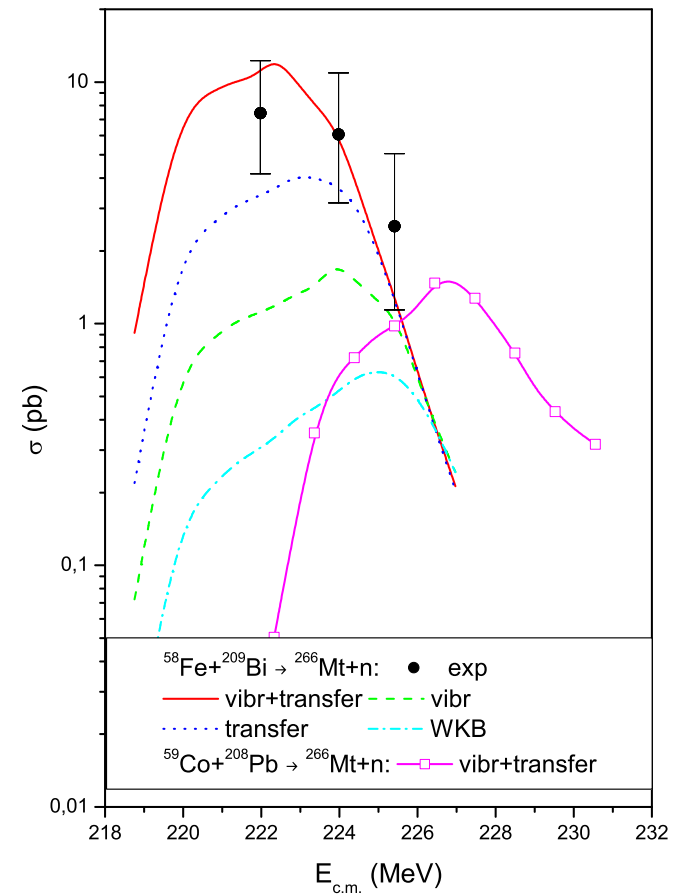
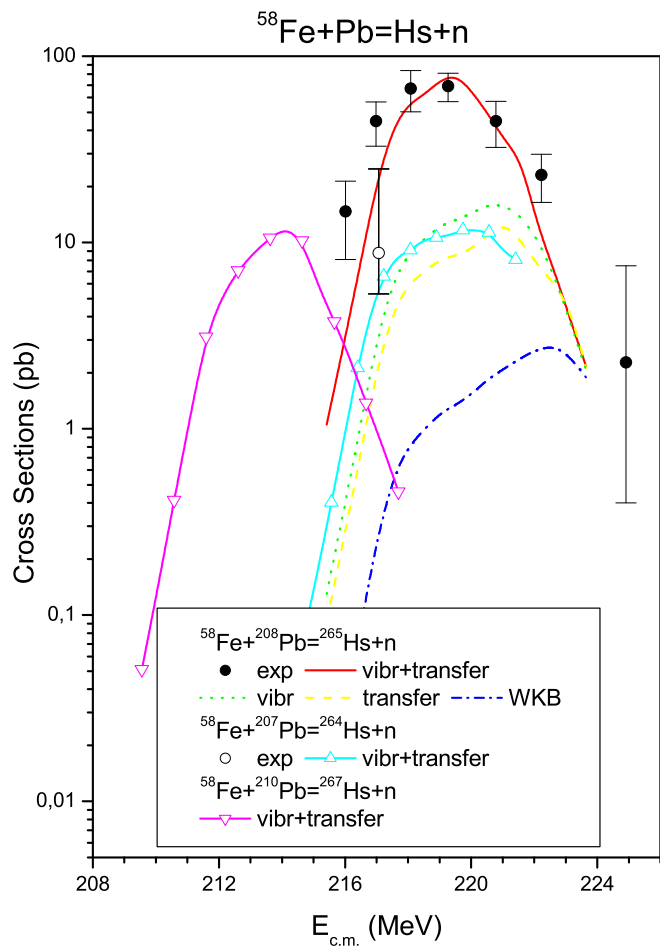
$$V_{\text{int}}(R, \beta_{21}, \beta_{22}) = V_{\text{Nucl}}(R, \beta_{21}, \beta_{22}) + V_{\text{Coul}}(R, \beta_{21}, \beta_{22})$$



$$V_{\text{int}}(R, \beta_{21}, \beta_{22}) = V_{\text{int}}(R, \beta_{21}, \beta_{22}) - Q_{\text{Transfer}}$$







Semi-Microscopic Potential (SMP) between heavy nuclei

The interaction energy between spherical and axial-symmetric nuclei in the frozen density approximation is

$$V(R, \vartheta) = E_{12}(R, \vartheta) - E_1 - E_2,$$

R is the distance between mass centers of colliding nuclei,

ϑ is the angle between the axial-symmetry axis of deformed nuclei and the line connected the mass centers of nuclei,

binding energies are

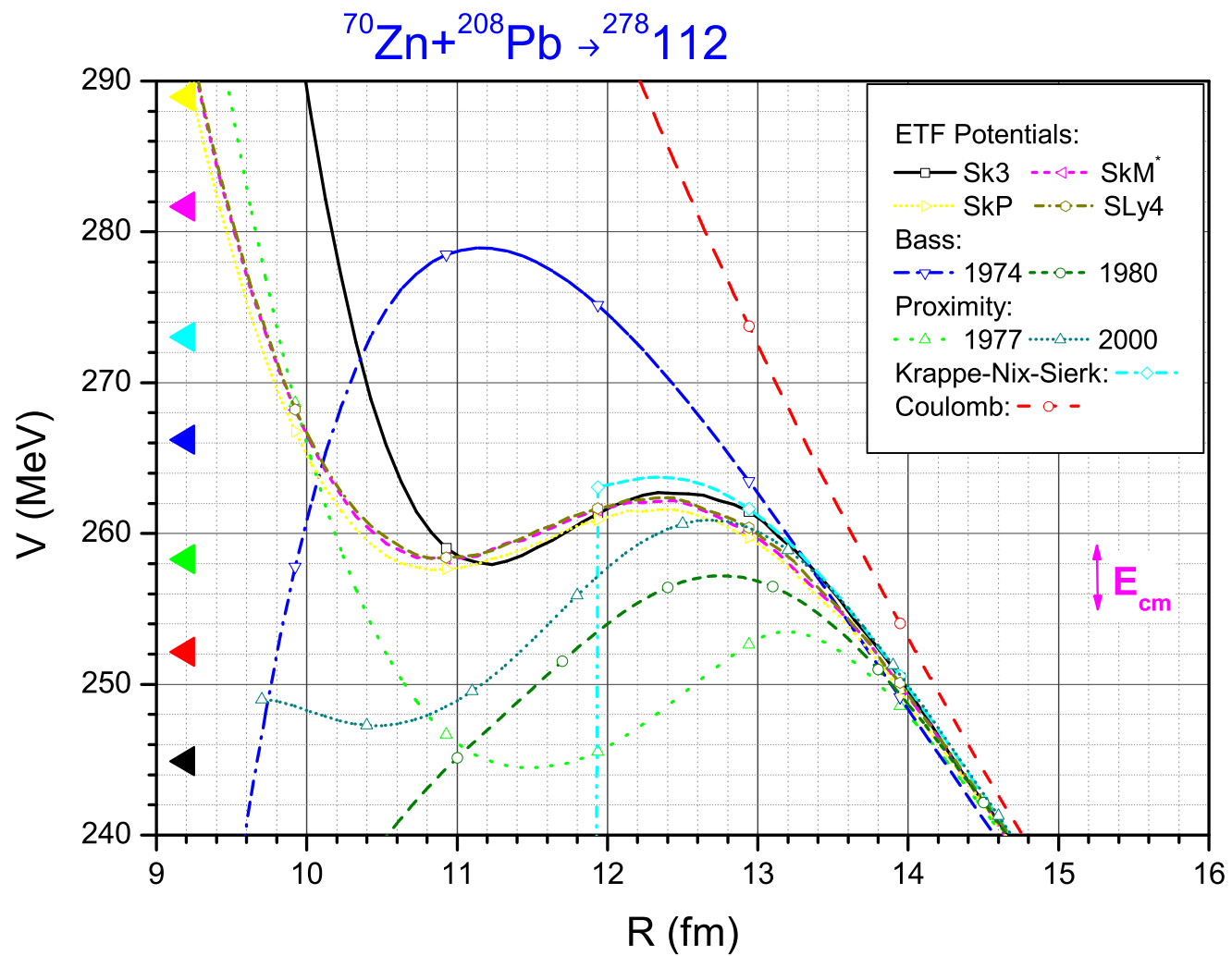
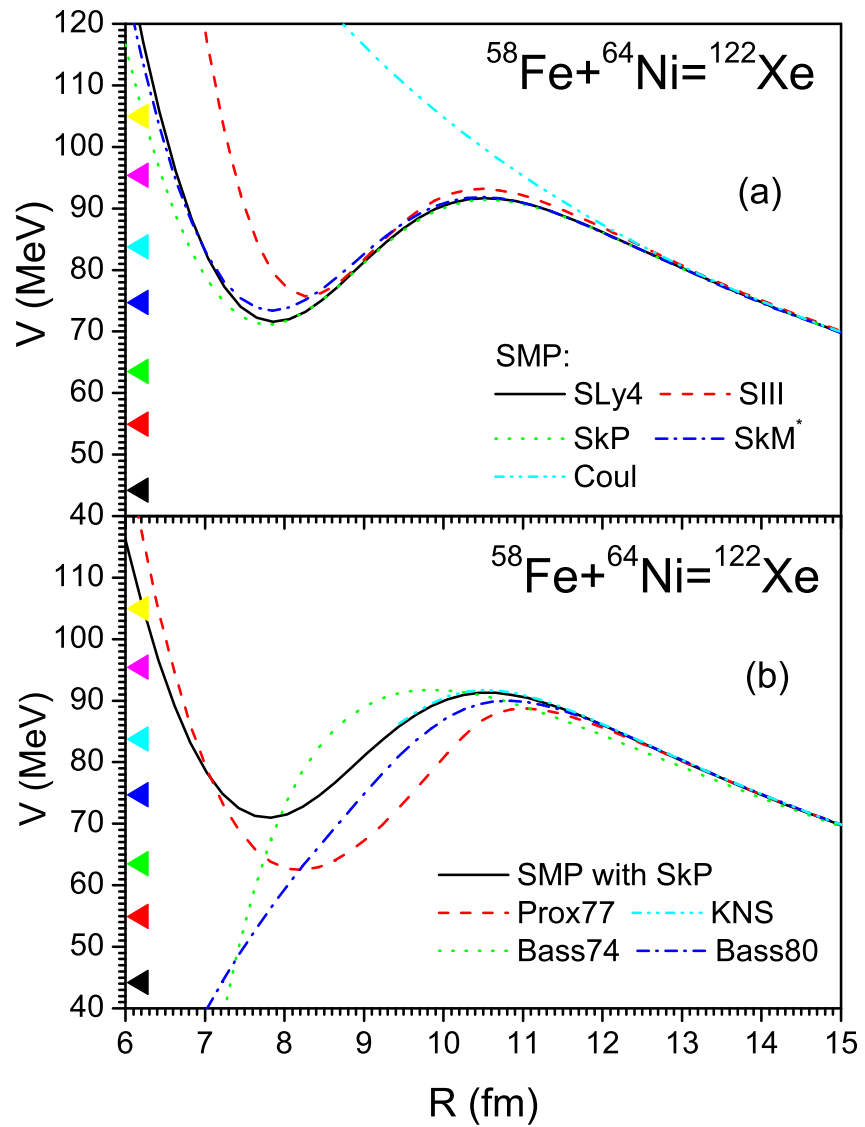
$$E_{12}(R, \vartheta) = \int \mathcal{E}[\rho_{1p}(\mathbf{r}) + \rho_{2p}(R, \vartheta, \mathbf{r}), \rho_{1n}(\mathbf{r}) + \rho_{2n}(R, \vartheta, \mathbf{r})] d\mathbf{r},$$
$$E_1 = \int \mathcal{E}[\rho_{1p}(\mathbf{r}), \rho_{1n}(\mathbf{r})] d\mathbf{r}, \quad E_2 = \int \mathcal{E}[\rho_{2p}(\mathbf{r}), \rho_{2n}(\mathbf{r})] d\mathbf{r}.$$

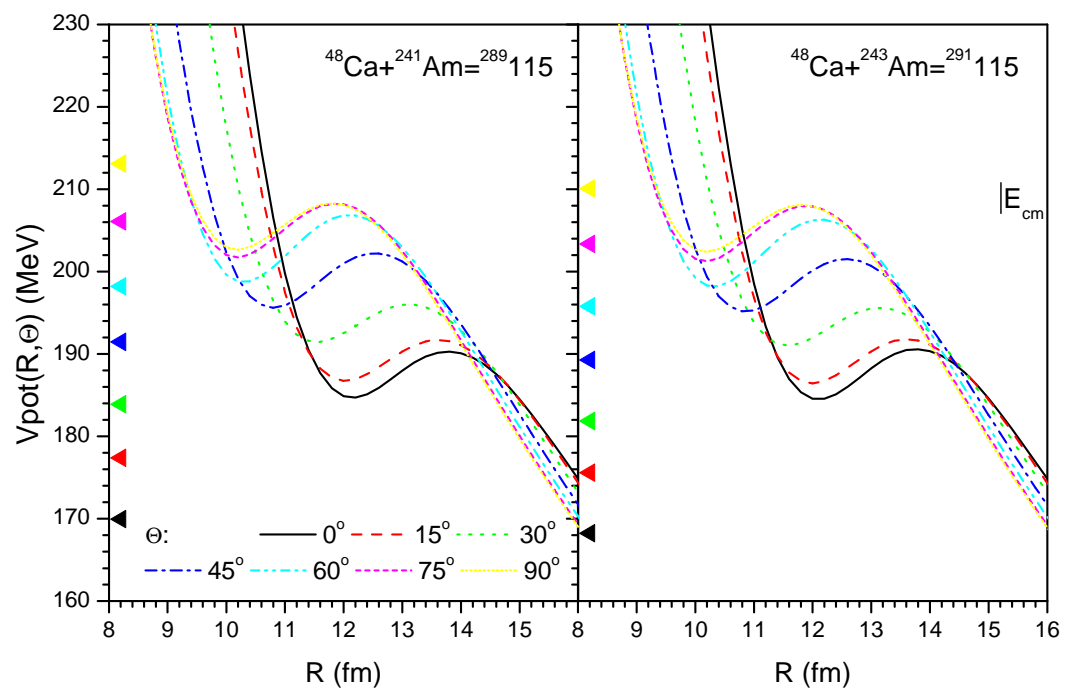
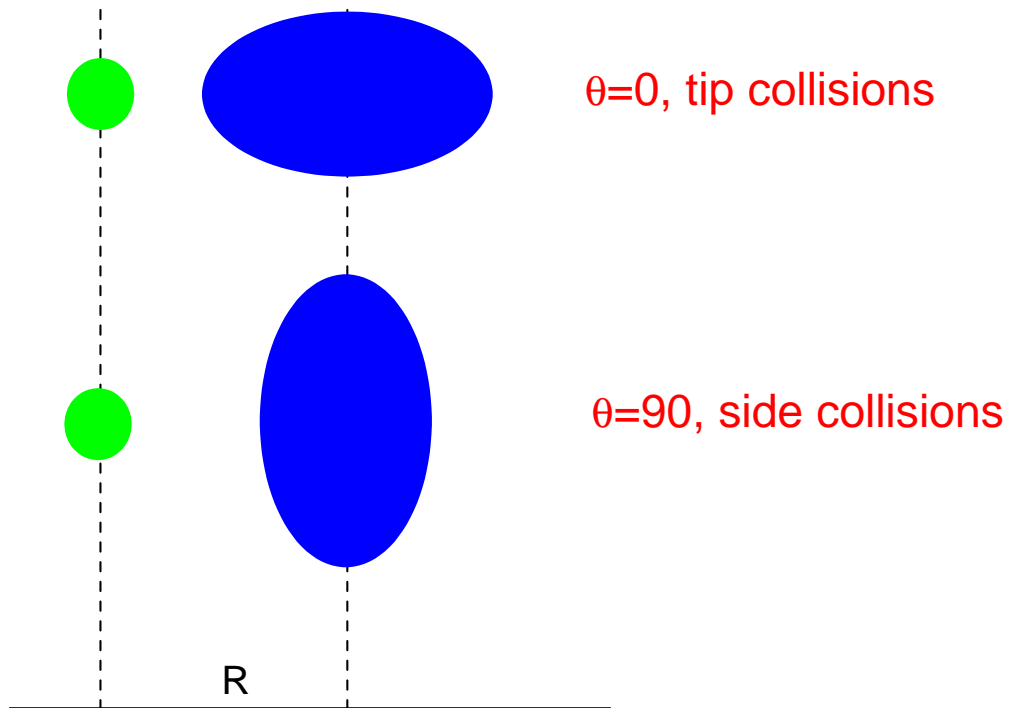
The energy density functional

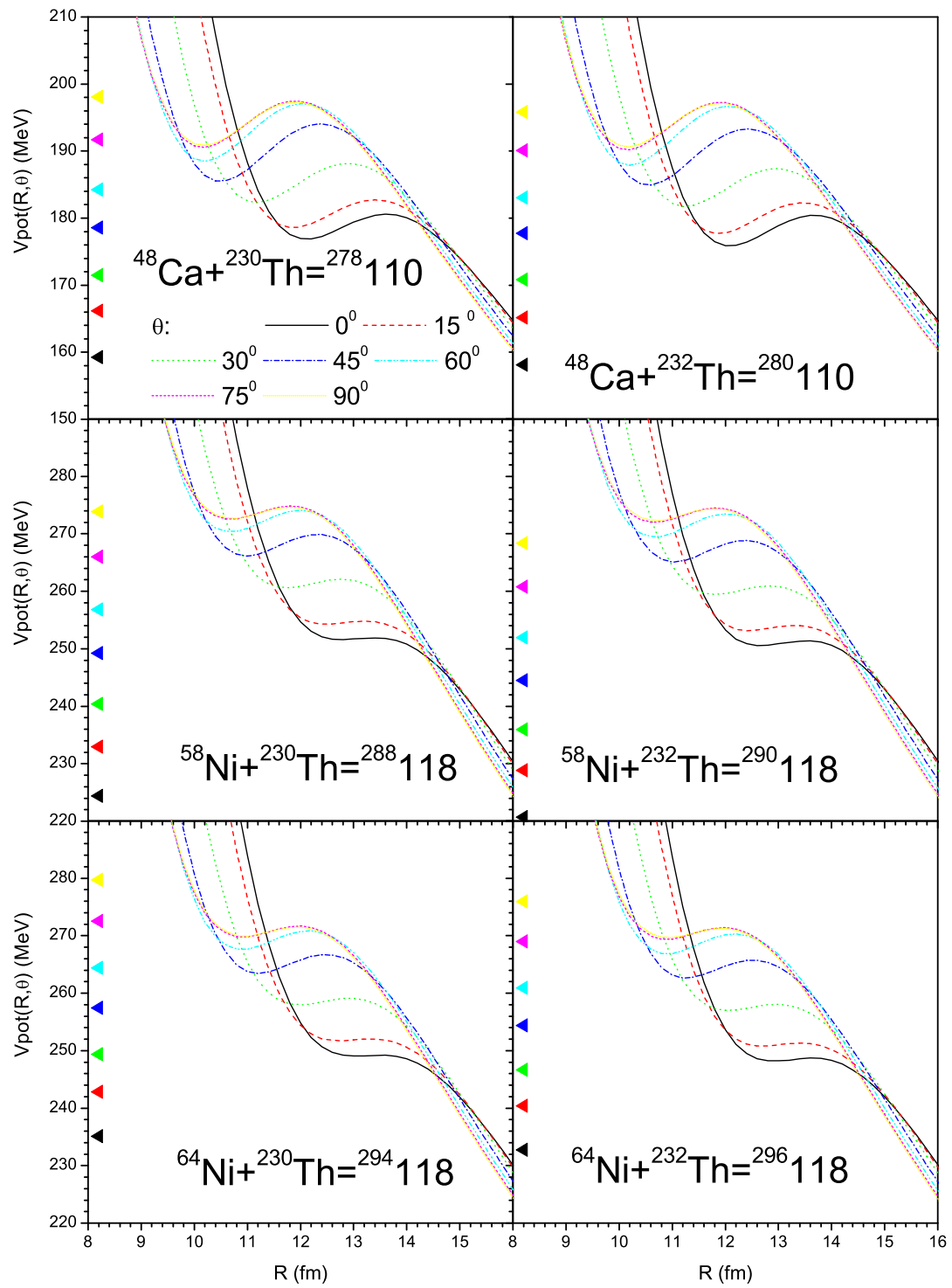
$$\mathcal{E}[\rho_p(\mathbf{r}), \rho_n(\mathbf{r})] = \frac{\hbar^2}{2m}[\tau_p(\mathbf{r}) + \tau_n(\mathbf{r})] + \mathcal{V}(\mathbf{r}),$$

where m is the nucleon mass. The expressions for proton τ_p and neutron τ_n kinetic energy density functionals are taken into account \hbar^2 corrections. The potential energy density functional splits into Skyrme and Coulomb (direct and exchange) parts

$$\mathcal{V}(\mathbf{r}) = \mathcal{V}_{\text{Skyrme}}(\mathbf{r}) + \mathcal{V}_{\text{Coul}}(\mathbf{r}).$$







The Coulomb interaction of two axial-symmetric nuclei

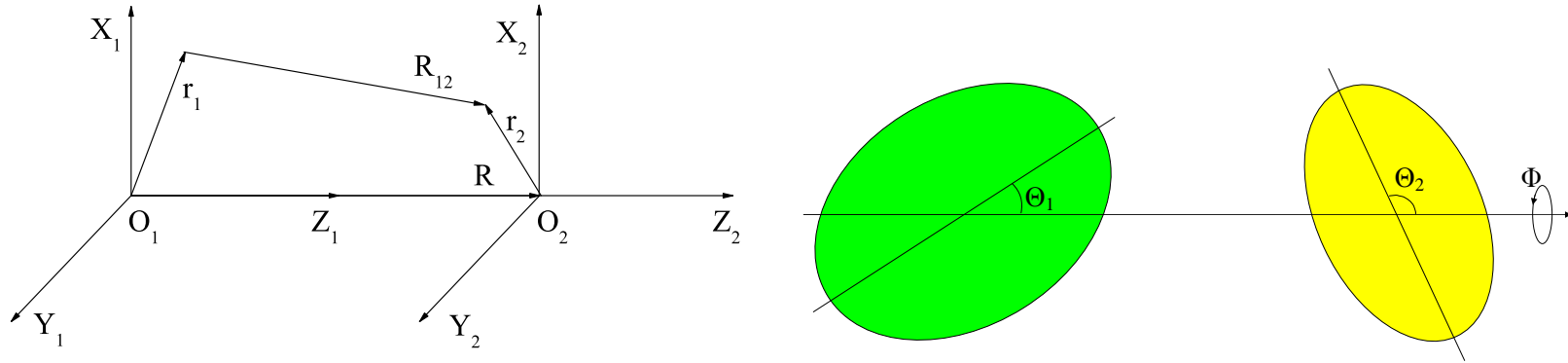
The Coulomb interaction of two nuclei at distances between their mass centers R is

$$V_C(R) = e^2 \int \frac{\rho_1(\mathbf{r}_1)\rho_2(\mathbf{r}_2)}{|\mathbf{R} + \mathbf{r}_2 - \mathbf{r}_1|} d\mathbf{r}_1 d\mathbf{r}_2.$$

where the denominator in can be presented as

$$\frac{1}{|\mathbf{R} + \mathbf{r}_2 - \mathbf{r}_1|} = \sum_{\ell_1, \ell_2=0}^{\infty} \frac{r_1^{\ell_1} r_2^{\ell_2}}{R^{\ell_1 + \ell_2 + 1}} \frac{4\pi(-1)^{\ell_2}(\ell_1 + \ell_2)!}{\sqrt{(2\ell_1 + 1)(2\ell_2 + 1)}} \times \sum_m \frac{Y_{\ell_1 m}(\vartheta_1, \varphi_1) Y_{\ell_2 - m}(\vartheta_2, \varphi_2)}{\sqrt{(\ell_1 + m)!(\ell_1 - m)!(\ell_2 + m)!(\ell_2 - m)!}},$$

where $Y_{\ell m}(\vartheta, \varphi)$ is the spherical harmonic functions, $r_i, \vartheta_i, \varphi_i$ are the spherical coordinates in the laboratory coordinate system O_i . Note that $Y_{\ell m}(\vartheta, \varphi)$ vanish whenever $|m| > \ell$.



The Coulomb interaction of two axial-symmetric arbitrary-oriented nuclei in the form

$$V_C(R, \Theta_1, \Theta_2, \Phi) = \frac{Z_1 Z_2 e^2}{R} \left\{ 1 + \sum_{\ell \geq 2} [f_{1\ell}(R, \Theta_1, R_{10})\beta_{1\ell} + f_{1\ell}(R, \Theta_2, R_{20})\beta_{2\ell}] \right. \\ \left. + f_2(R, \Theta_1, R_{10})\beta_{12}^2 + f_2(R, \Theta_2, R_{20})\beta_{22}^2 + f_3(R, \Theta_1, \Theta_2, R_{10}, R_{20})\beta_{12}\beta_{22} + f_4(R, \Theta_1, \Theta_2, \Phi, R_{10}, R_{20})\beta_{12}\beta_{22} \right\}.$$

We approximate the nuclear part of the potential between deformed nuclei as

$$V_n(R, \Theta_1, \Theta_2, \Phi) \approx \frac{C_{10} + C_{20}}{C_{\text{def}}} V_n^0(d^0(R_{\text{sph}}, R_{10}, R_{20})),$$

where

$$C_{\text{def}} = \left[(C_1^{\parallel} + C_2^{\parallel})(C_1^{\perp} + C_2^{\perp}) \right]^{1/2}$$

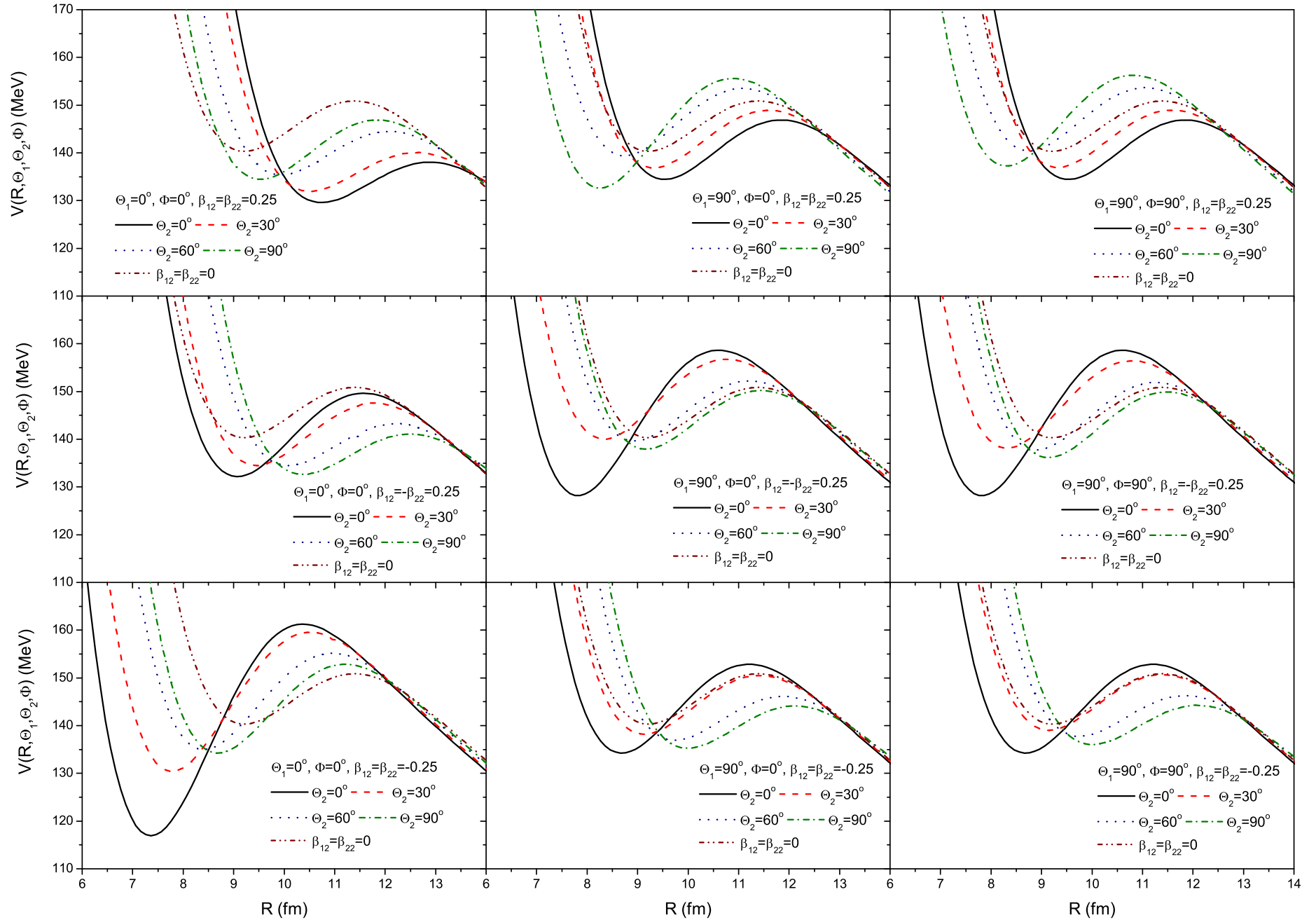
is the generalized reciprocal curvature and

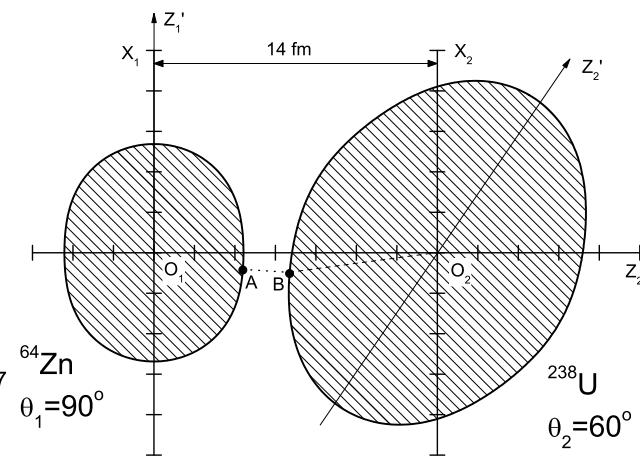
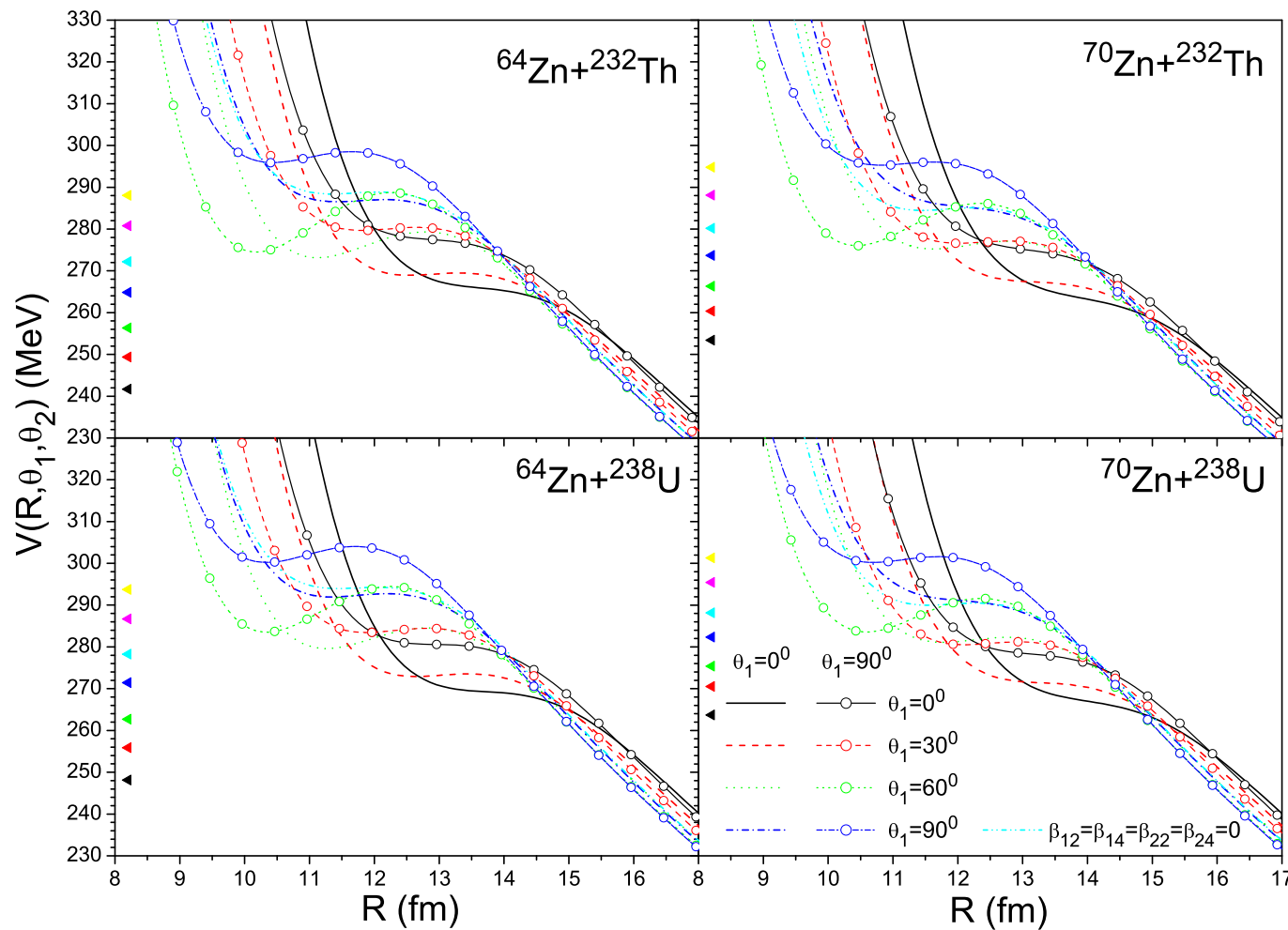
$$d^0(R_{\text{sph}}, R_{10}, R_{20}) = d(R, \Theta_1, \Theta_2, \Phi, R_{10}, R_{20}, \beta_1, \beta_2).$$

The nucleus-nucleus interaction potential of axial-symmetric nuclei at various relative orientations for zero value of orbital momentum is given as

$$V(R, \Theta_1, \Theta_2, \Phi) = V_C(R, \Theta_1, \Theta_2, \Phi) + V_n(R, \Theta_1, \Theta_2, \Phi).$$

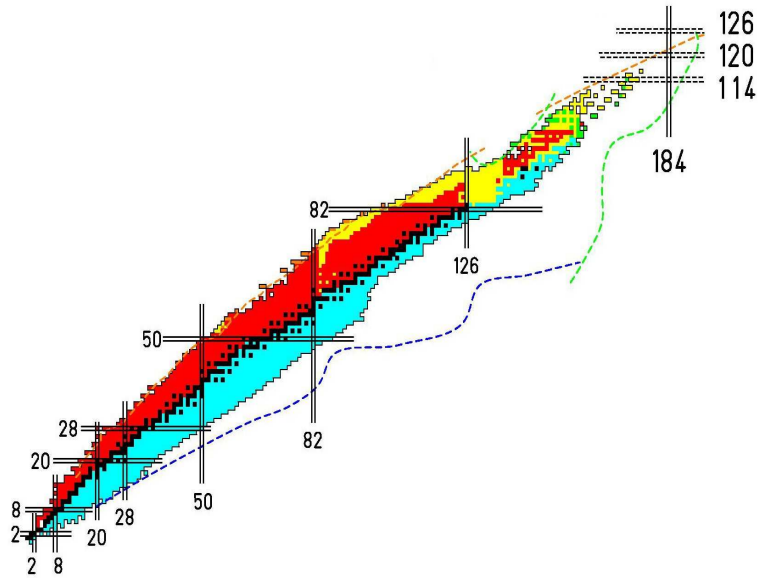
Potentials ${}^{86}\text{Kr}(\beta_{12} = \pm 0.25) + {}^{86}\text{Kr}(\beta_{22} = \pm 0.25)$





Magic numbers for ultraheavy region. Alpha-decay half-life

and fission barriers for double-magic ultraheavy nuclei



Magic Numbers $Z = 8, 20, 28, 50, 82, N = 8, 20, 28, 50, 82, 126$ are related to filling of the nucleonic shells.

Nuclei with magic number(s) of nucleons have:

⇒ higher stability

⇒ higher stiffness to perturbation of different nature

⇒ higher abundance in the nature

then neighboring nuclei.

The N or Z dependencies of nucleon(s) separation energy have jump at crossing magic numbers.

The largest magic numbers "experimentally known" up to now are: $Z = 114, N = 184$.

The P or N shell corrections have deep local minimum at magic numbers.

TASKs: 1. determinate magic numbers for heavy systems with $A = 300 - 1200$ by studying deep local minimum of shell corrections.

2. evaluate $T_{1/2}(\alpha)$ and fission barrier for magic ultraheavy.

Ultra-heavy nuclei can exist in star (neutron stars, magnetars, and etc)

Green's β -stability line:

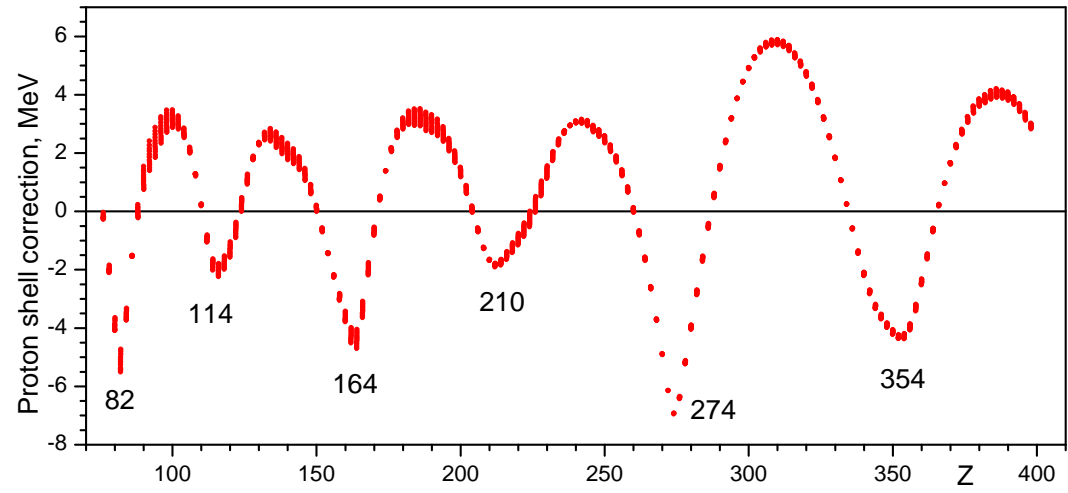
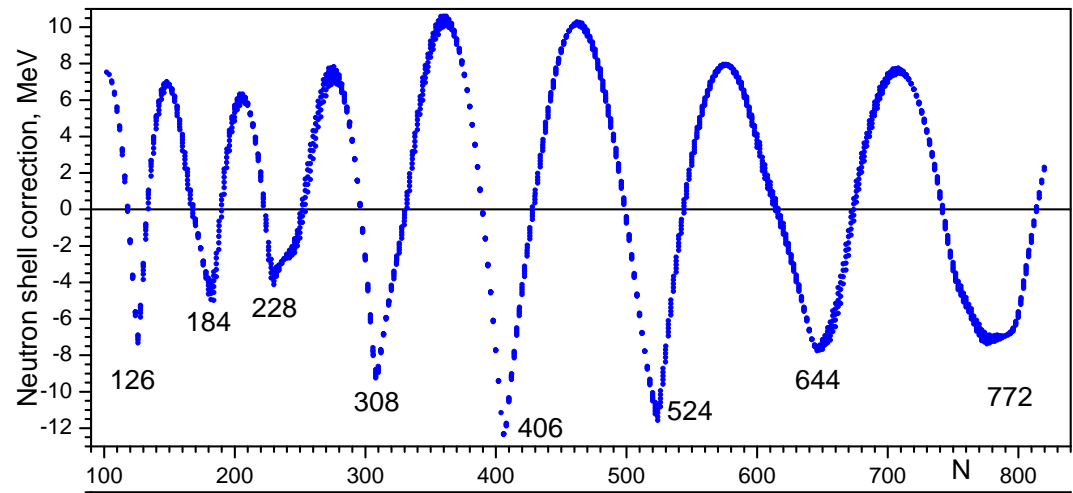
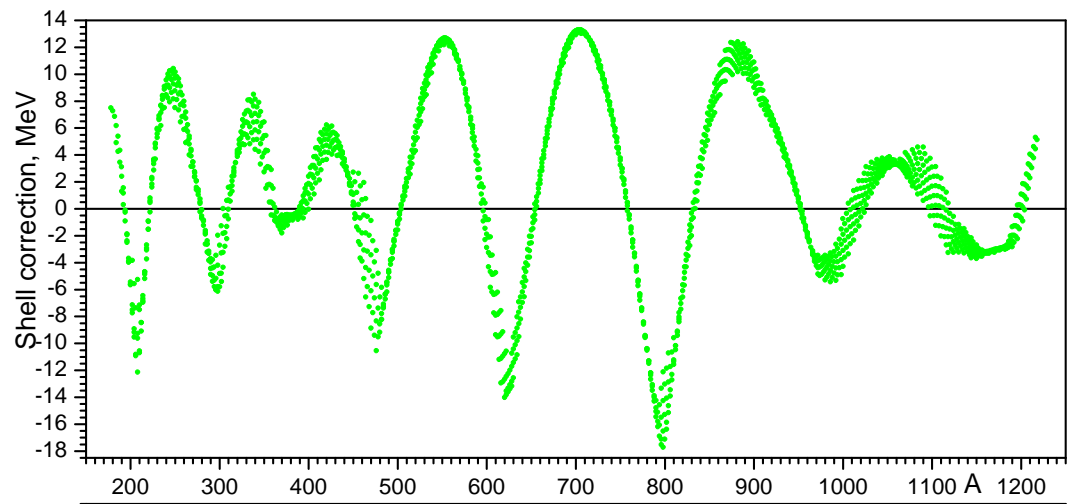
$$N_{Green} = \frac{2}{3}Z + \frac{5}{3}(10000 + 40Z + Z^2)^{1/2} - \frac{500}{3}. \quad (1)$$

We evaluate Proton δ_P and Neutron δ_N shell corrections for even-even nuclei along Green's β -stability line for $Z = 76 - 400$, $N = 102 - 820$, $A = 178 - 1218$.

$$Z = 76, 78, 80, \dots 400.$$

$$N = N_{Green} - 10, N_{Green} - 8, N_{Green} - 6, \dots N_{Green} + 10$$

Mean field potential: Woods-Saxon + spin-orbit interaction + Lipkin-Nogami pairing + Coulomb (protons). Parametrization : "universal" - S.Cwiok, et al, CPC **46** 379 (1987).



WE OBTAINED: Proton Magic Numbers: $Z = 82, 114, 164, 210, 274, 354$
Neutron Magic Numbers: $N = 126, 184, 228, 308, 406, 524, 644, 772$

Comparison with other calculations:

$Z=82, 114, 164$ and $N=126, 184, 228 \Leftrightarrow$ Shell model.

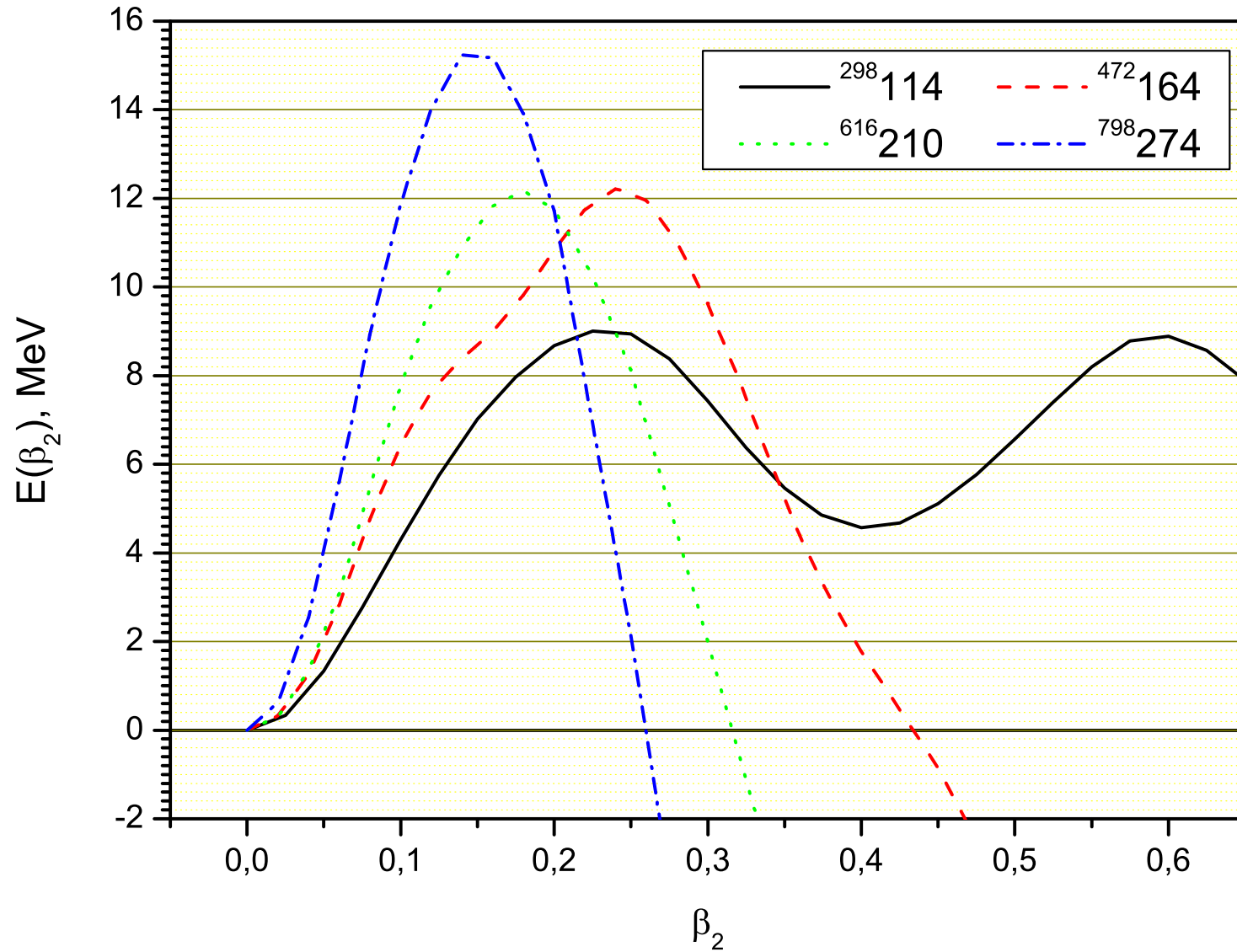
S.G. Nilsson, I. Ragnarsson I., Shapes and shells in nuclear structure (Cambridge Univ. Press, Cambridge, 1995).

$Z \approx 114-126, Z \approx 164, N \approx 172-184, N \approx 228, N \approx 308, N \approx 406 \Leftrightarrow$ HFB+Skyrme or RMF

There is strong dependence on the both model and force set.

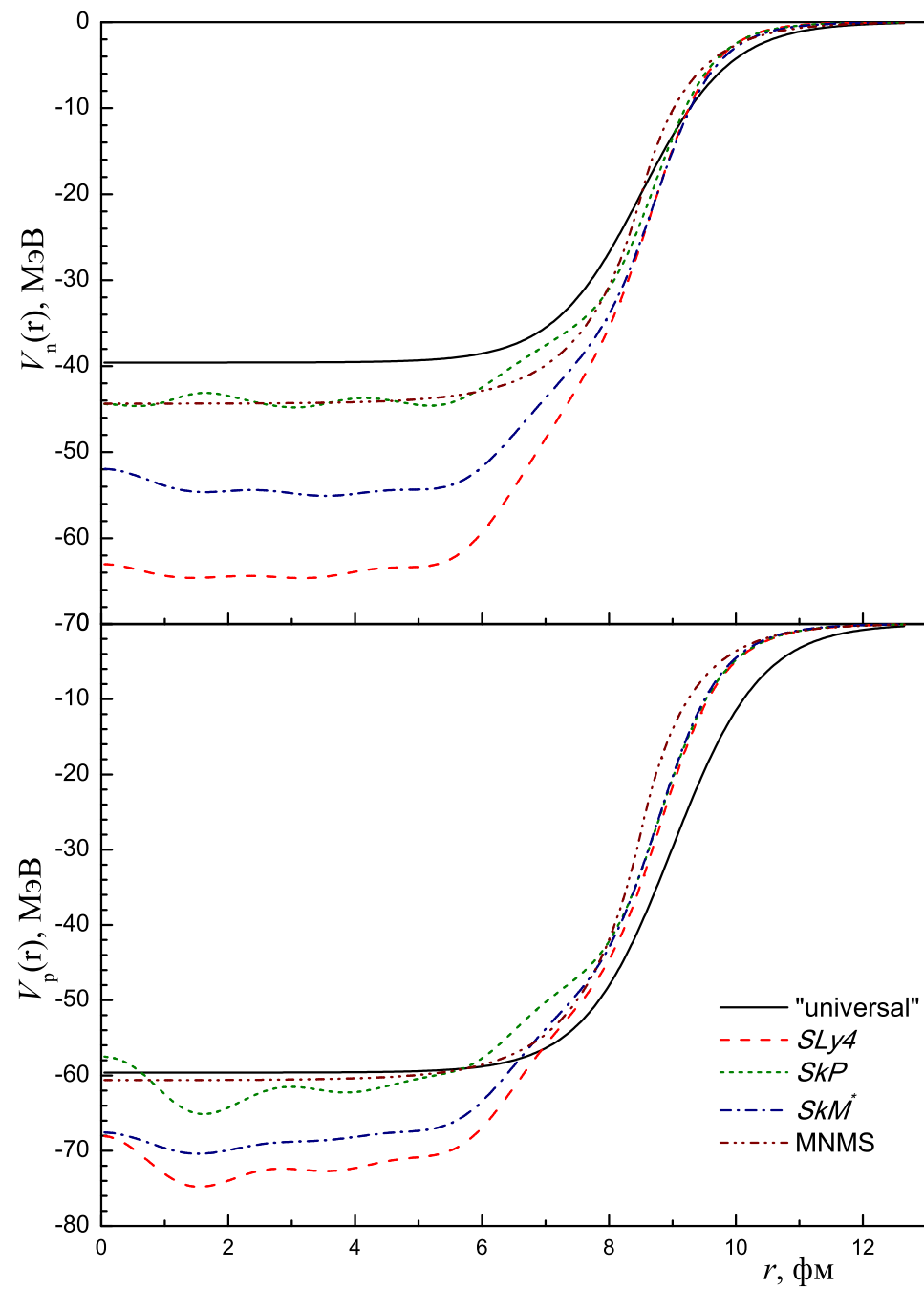
M. Bender, W. Nazarewicz, P.-G. Reinhard, Phys. Lett. **B515**, 42 (2001).

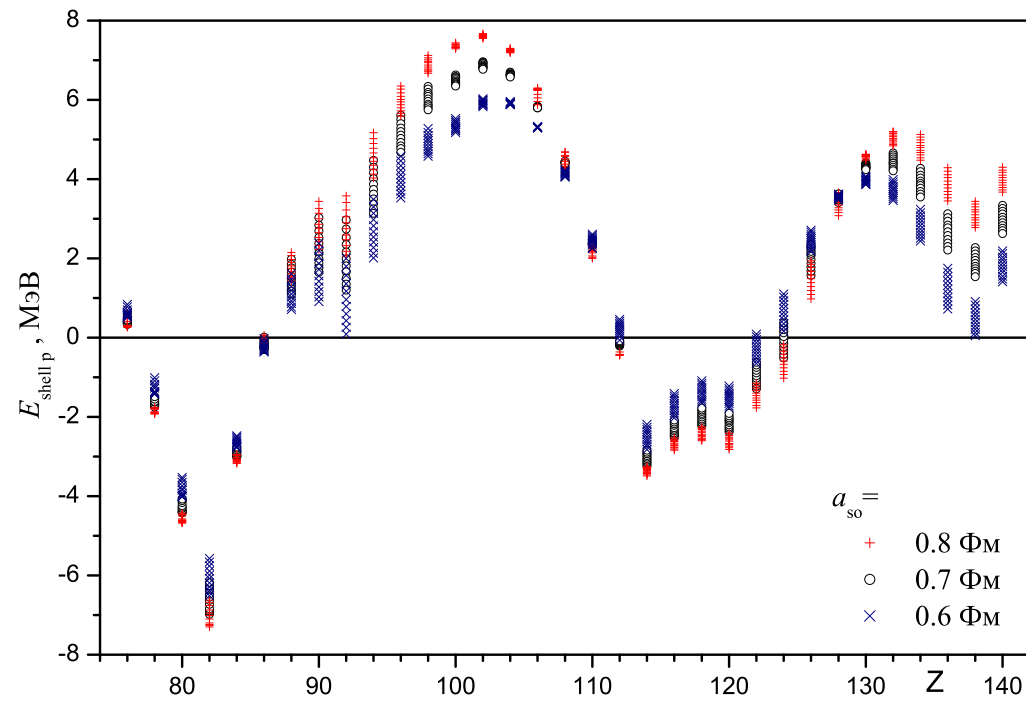
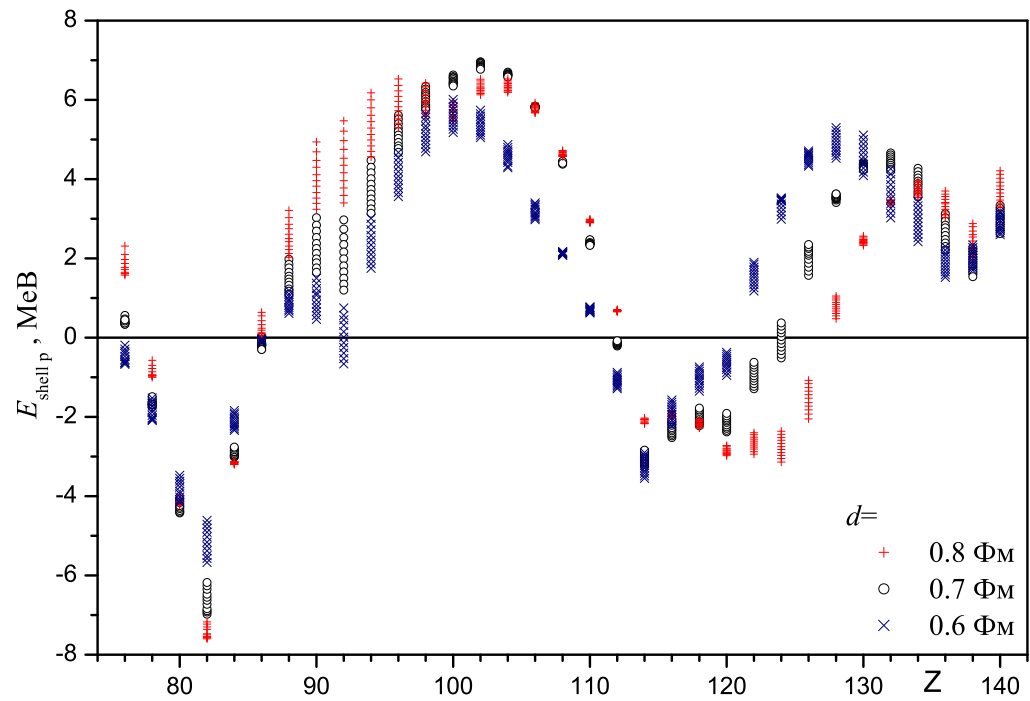
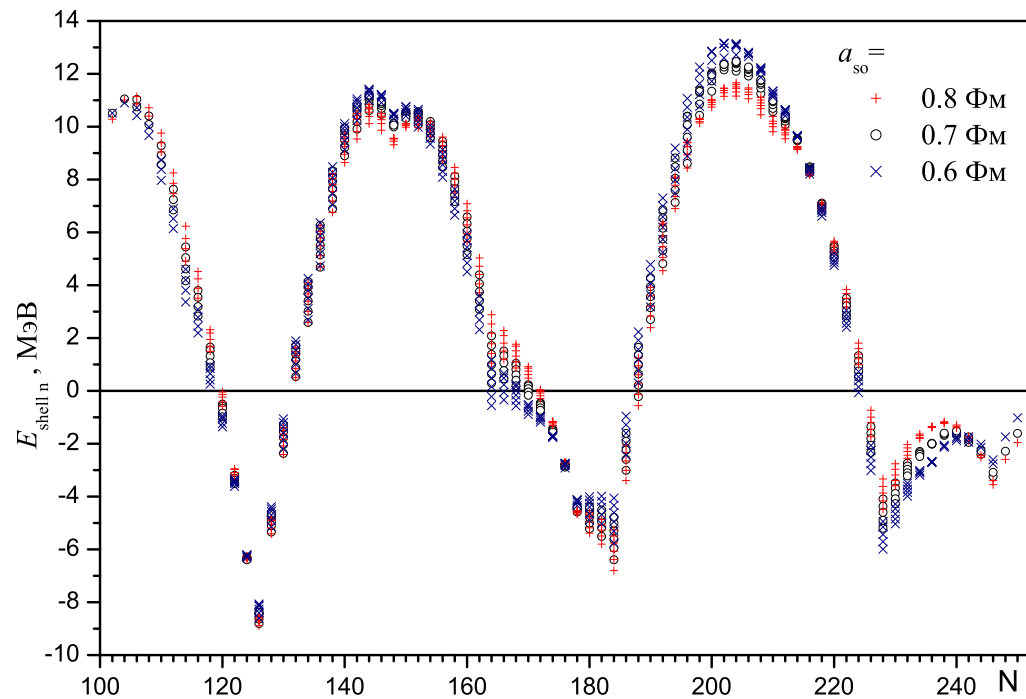
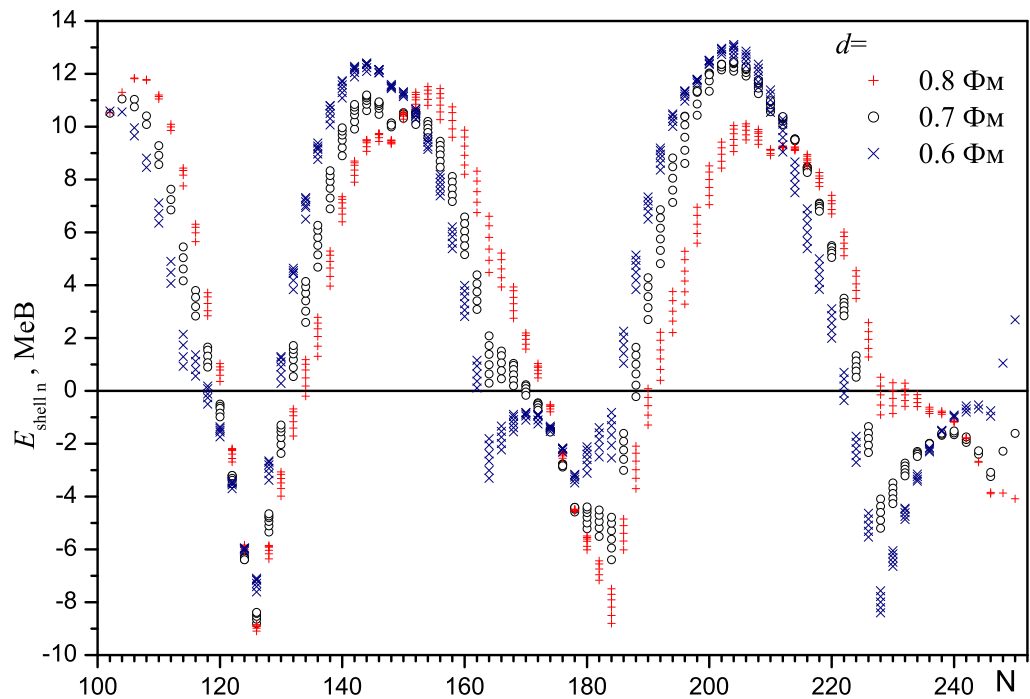
Fission barriers of ultraheavy double-magic nuclei.



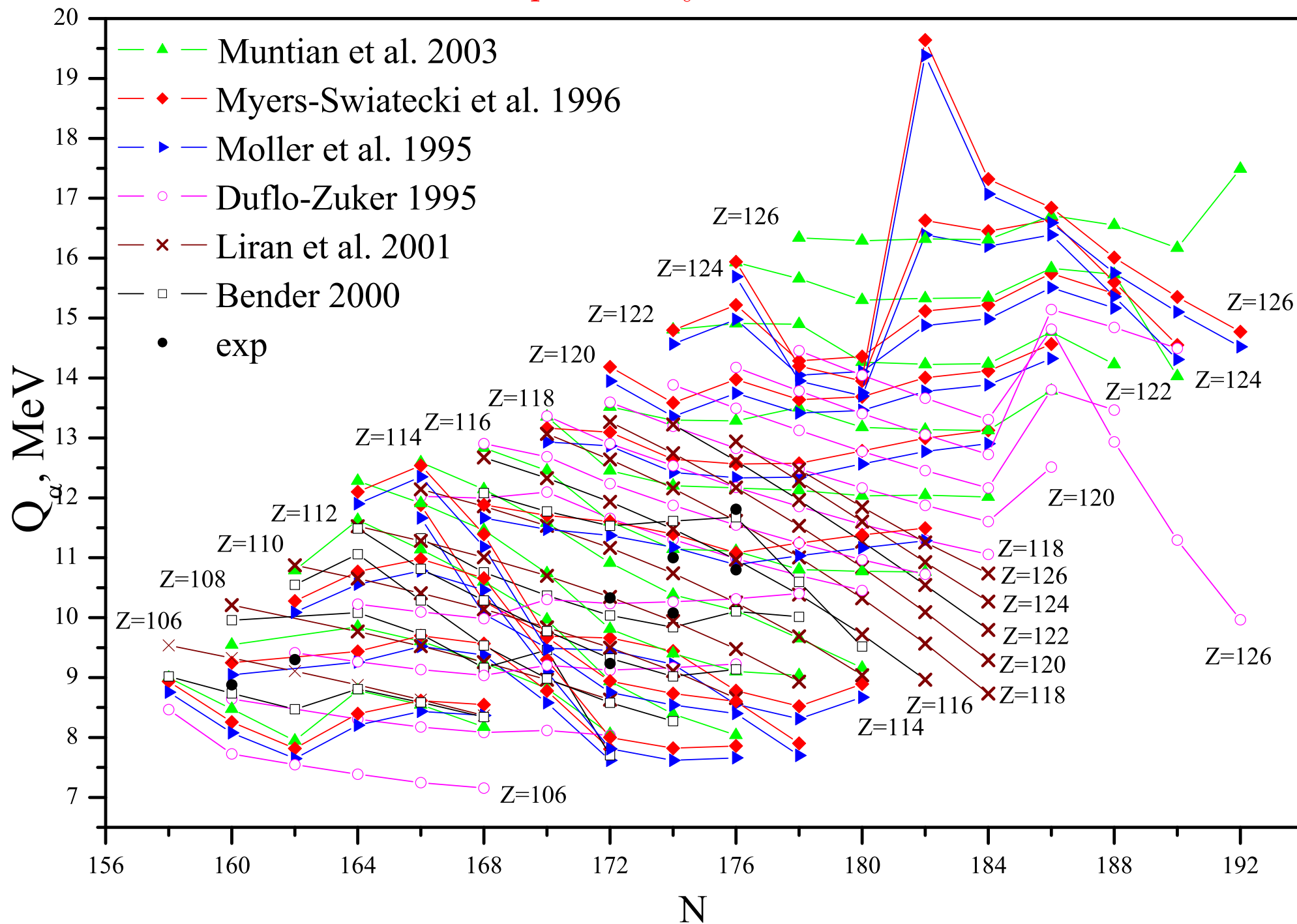
Deformation energy \Leftrightarrow Droplet Model (Yukawa+exp) + shell correction

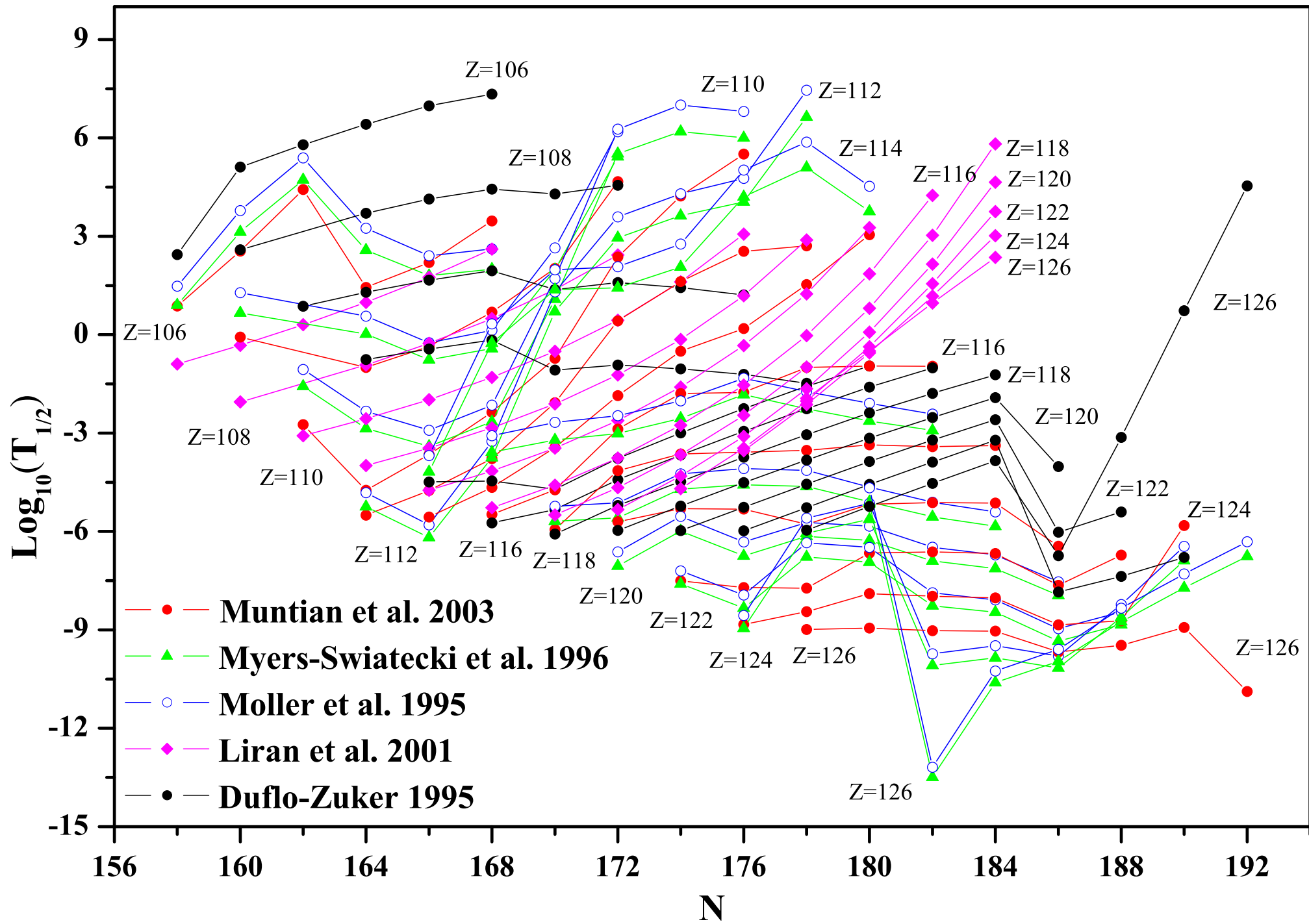
Dependence of magic number in SHE region on mean-field approximation





Alpha-decay of SHE





Thanks for your attention!